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Physics 171

Solution Set #3

Fall 2001

②

① (a) The simplest way to verify that raising and lowering indices commutes with covariant differentiation is to note that

$$\begin{aligned} A_{\mu;\alpha} &= (g_{\mu\nu} A^\nu)_{;\alpha} \\ &= g_{\mu\nu;\alpha} A^\nu + g_{\mu\nu} A^{\nu;\alpha} \end{aligned}$$

using Leibniz's rule for differentiation of a product. But

$$g_{\mu\nu;\alpha} = 0$$

Hence,
$$A_{\mu;\alpha} = g_{\mu\nu} A^{\nu;\alpha}$$

The same argument works for

$$A^{\mu;\alpha} = (g^{\mu\nu} A_\nu)_{;\alpha} = g^{\mu\nu} A_{\nu;\alpha}$$

since $g^{\mu\nu}_{;\alpha} = 0$. The argument is easily extended to raising or lowering the index of any tensor.

(b) Following the examples given in class,

$$A^\alpha{}_{\beta\beta;\sigma} = A^\alpha{}_{\beta\beta,\sigma} + \Gamma^\alpha{}_{\mu\sigma} A^\mu{}_{\beta\beta} - \Gamma^\mu{}_{\beta\sigma} A^\alpha{}_{\mu\beta} - \Gamma^\mu{}_{\beta\sigma} A^\alpha{}_{\beta\mu}$$

(c) By definition, the metric connection is

$$\Gamma^\mu{}_{\alpha\beta} = \frac{1}{2} g^{\mu\nu} [g_{\nu\beta;\alpha} + g_{\nu\alpha;\beta} - g_{\alpha\beta;\nu}]$$

If μ, α and β are all different, and if $g_{\nu\beta}$ is diagonal then $g_{\nu\beta} = 0$. Moreover, in the implicit sum over ν , we must have $\mu = \nu$ otherwise $g^{\mu\nu} = 0$. Note that $g^{\mu\nu}$ is the inverse of $g_{\mu\nu}$ and is therefore diagonal as well. Finally, since $\mu = \nu$ but ν differs from α and β , we see that $g_{\nu\beta} = g_{\nu\alpha} = 0$. We conclude that if $g_{\mu\nu}$ is diagonal and μ, α , and β are distinct, then $\Gamma^\mu{}_{\alpha\beta} = 0$.

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② (a)
$$A_{\mu;\nu} = A_{\mu,\nu} - A_\alpha \Gamma^\alpha{}_{\mu\nu}$$

$$\begin{aligned} A_{\nu;\mu} &= A_{\nu,\mu} - A_\alpha \Gamma^\alpha{}_{\nu\mu} \\ &= A_{\nu,\mu} - A_\alpha \Gamma^\alpha{}_{\mu\nu} \end{aligned}$$

since $\Gamma^\alpha{}_{\mu\nu} = \Gamma^\alpha{}_{\nu\mu}$. Hence

$$A_{\mu;\nu} - A_{\nu;\mu} = A_{\mu,\nu} - A_{\nu,\mu}$$

(b)
$$F_{\mu\nu;\rho} = F_{\mu\nu,\rho} - \Gamma^\sigma{}_{\mu\rho} F_{\sigma\nu} - \Gamma^\sigma{}_{\nu\rho} F_{\mu\sigma}$$

$$F_{\rho\mu;\nu} = F_{\rho\mu,\nu} - \Gamma^\sigma{}_{\rho\nu} F_{\sigma\mu} - \Gamma^\sigma{}_{\mu\nu} F_{\rho\sigma}$$

$$F_{\nu\rho;\mu} = F_{\nu\rho,\mu} - \Gamma^\sigma{}_{\nu\mu} F_{\sigma\rho} - \Gamma^\sigma{}_{\rho\mu} F_{\nu\sigma}$$

Adding the three equations and using the fact that

$$F_{\mu\sigma} = -F_{\sigma\mu}$$

$$\Gamma^\sigma{}_{\rho\nu} = \Gamma^\sigma{}_{\nu\rho}$$

we see that

$$\begin{aligned} F_{\mu\nu;\rho} + F_{\rho\mu;\nu} + F_{\nu\rho;\mu} &= F_{\mu\nu,\rho} + F_{\rho\mu,\nu} + F_{\nu\rho,\mu} \\ &\quad - \Gamma^\sigma{}_{\mu\rho} (F_{\sigma\nu} + F_{\nu\sigma}) - \Gamma^\sigma{}_{\nu\rho} (F_{\mu\sigma} + F_{\sigma\mu}) - \Gamma^\sigma{}_{\rho\mu} (F_{\nu\sigma} + F_{\sigma\nu}) \end{aligned}$$

or

$$F_{\mu\nu;\rho} + F_{\rho\mu;\nu} + F_{\nu\rho;\mu} = F_{\mu\nu,\rho} + F_{\rho\mu,\nu} + F_{\nu\rho,\mu}$$

(c) In Minkowski space, Maxwell's equations are:

$$\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} J^\nu$$

$$\partial^\rho F^{\mu\nu} + \partial^\mu F^{\nu\rho} + \partial^\nu F^{\rho\mu} = 0$$

where
$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

If we lower the indices, we can write

$$\partial_\rho F_{\mu\nu} + \partial_\mu F_{\nu\rho} + \partial_\nu F_{\rho\mu} = 0$$

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$$\text{where } F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu.$$

Using the comma notation, we therefore have:

$$F^{\mu\nu}_{;\mu} = \frac{4\pi}{c} J^\nu$$

$$F_{\mu\nu;\rho} + F_{\rho\mu;\nu} + F_{\rho\nu;\mu} = 0$$

$$\text{where } F_{\mu\nu} = A_{\nu;\mu} - A_{\mu;\nu}.$$

To extend Maxwell's equations to curved spacetime, we make the above equations generally covariant. To accomplish this, we replace ordinary derivatives with covariant derivatives. For example, we now define

$$F_{\mu\nu} = A_{\nu;\mu} - A_{\mu;\nu}$$

Note that based on part (a), we can simply write this as

$$F_{\mu\nu} = A_{\nu;\mu} - A_{\mu;\nu}$$

so there is no change in the definition of $F_{\mu\nu}$. But now,

$$F^{\mu\nu} = g^{\alpha\mu} g^{\nu\beta} F_{\alpha\beta}$$

where $g^{\mu\alpha}$ is the inverse of the metric tensor of curved spacetime. The Maxwell equations in covariant form are now:

$$F^{\mu\nu}_{;\mu} = \frac{4\pi}{c} J^\nu$$

$$F_{\mu\nu;\rho} + F_{\rho\mu;\nu} + F_{\rho\nu;\mu} = 0$$

Based on the results of part (b), the second equation can be rewritten as

$$F_{\mu\nu;\rho} + F_{\rho\nu;\mu} + F_{\rho\mu;\nu} = 0$$

so again, there is no change from the Minkowski space version.

Finally, current conservation in Minkowski space is $\partial_\mu J^\mu = 0$, which we can write as $J^{\mu}_{;\mu} = 0$. In curved spacetime, we again replace the ordinary derivative with a covariant derivative. Thus, in curved spacetime,

$$J^{\mu}_{;\mu} = 0.$$

Remark: If one takes the covariant derivative of $F^{\mu\nu}_{;\mu} = \frac{4\pi}{c} J^\nu$ with respect to x^ν , one can verify after much algebra that

$$(F^{\mu\nu}_{;\mu})_{;\nu} = 0$$

In deriving this result, one must use the fact that $F^{\mu\nu} = -F^{\nu\mu}$. Thus, Maxwell's equation is consistent with the requirement that $J^{\mu}_{;\mu} = 0$.

④

③ (a) Start with the result:

$$T^{\alpha}_{\mu\nu} = \frac{1}{2} g^{\alpha\rho} [g_{\rho\nu;\mu} + g_{\rho\mu;\nu} - g_{\mu\nu;\rho}]$$

Setting $\alpha = \mu$ and summing over μ ,

$$T^{\mu}_{\mu\nu} = \frac{1}{2} g^{\mu\rho} [g_{\rho\nu;\mu} + g_{\rho\mu;\nu} - g_{\mu\nu;\rho}]$$

Note that $g^{\mu\rho} [g_{\rho\nu;\mu} - g_{\mu\nu;\rho}] = 0$ since $g^{\mu\rho} = g^{\rho\mu}$ while $g_{\rho\nu;\mu} - g_{\mu\nu;\rho}$ is antisymmetric under the interchange of μ and ρ .

[Equivalently, one can simply relabel the second term $g^{\mu\rho} g_{\mu\nu;\rho}$ by writing ρ for μ and μ for ρ . This is allowed since ρ and μ are dummy indices; that is they are summed indices.]

Hence,

$$\begin{aligned} T^{\mu}_{\mu\nu} &= \frac{1}{2} g^{\mu\rho} g_{\rho\mu;\nu} \\ &= \frac{1}{2} g^{\mu\rho} \frac{\partial g_{\rho\mu}}{\partial x^\nu} \end{aligned}$$

To prove that this is equivalent to

$$T^{\mu}_{\mu\nu} = \frac{1}{2g} \frac{\partial g}{\partial x^\nu}, \quad g \equiv \det g_{\mu\nu}$$

Let us consider an arbitrary matrix A . By definition of the differential,

$$d(\det A) = \det(A+dA) - \det A$$

We can perform the following manipulation:

$$\begin{aligned} \det(A+dA) &= \det(A[1 + A^{-1}dA]) \\ &= \det A \det(1 + A^{-1}dA) \end{aligned}$$

where 1 is the identity matrix (with matrix elements δ_{ij}) and we have used the fact that $\det AB = \det A \det B$.

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I can evaluate $\det(1 + \delta A)$ as follows:

$$\det(1 + \delta A) = \det \begin{pmatrix} 1 + \delta A_{11} & \delta A_{12} & \dots & \delta A_{1n} \\ \delta A_{21} & 1 + \delta A_{22} & \dots & \delta A_{2n} \\ \vdots & \vdots & \ddots & \vdots \\ \delta A_{n1} & \delta A_{n2} & \dots & 1 + \delta A_{nn} \end{pmatrix}$$

$$= (1 + \delta A_{11})(1 + \delta A_{22}) \dots (1 + \delta A_{nn}) + O((\delta A)^n)$$

$$= 1 + \delta A_{11} + \delta A_{22} + \dots + \delta A_{nn} + O((\delta A)^2)$$

$$= 1 + \text{Tr}(\delta A)$$

where I have dropped all higher order corrections in δA_{ij} . Note that Tr stands for trace of the matrix, which is equal to the sum of the diagonal elements.

$$\text{Taking } \delta A = A^{-1} dA$$

it follows that to leading order in dA ,

$$\det(A + dA) = \det A [1 + \text{Tr}(A^{-1} dA)]$$

$$= \det A + \det A \text{Tr}(A^{-1} dA)$$

Hence,

$$d(\det A) = \det A \text{Tr}(A^{-1} dA)$$

which is true for any matrix A . Applying this result to $A = g_{\mu\nu}$, and noting that by definition, $A^{-1} = g^{\mu\nu}$, it follows that

$$\text{Tr}(A^{-1} dA) = g^{\mu\nu} dg_{\mu\nu}$$

$$= g^{\mu\nu} dg_{\mu\nu}$$

since $g^{\mu\nu} = g^{\nu\mu}$. Moreover, $\det A = g$. Hence

$$dg = g g^{\mu\nu} dg_{\mu\nu}$$

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Finally, using the chain rule,

$$dg = \frac{\partial g}{\partial x^\beta} dx^\beta$$

$$dg_{\mu\nu} = \frac{\partial g_{\mu\nu}}{\partial x^\beta} dx^\beta$$

$$\text{Hence, } \frac{1}{g} \frac{\partial g}{\partial x^\beta} dx^\beta = g^{\mu\nu} \frac{\partial g_{\mu\nu}}{\partial x^\beta} dx^\beta$$

This is true for an arbitrary infinitesimal dx^β , so we conclude that:

$$\frac{1}{g} \frac{\partial g}{\partial x^\beta} = g^{\mu\nu} \frac{\partial g_{\mu\nu}}{\partial x^\beta}$$

$$\text{Comparing with the previous result, } \Gamma^{\mu}_{\nu\mu} = \frac{1}{2} g^{\mu\sigma} \frac{\partial g_{\sigma\mu}}{\partial x^\nu}$$

(and using $g^{\mu\nu} = g^{\nu\mu}$), we see that

$$\Gamma^{\mu}_{\nu\mu} = \frac{1}{2g} \frac{\partial g}{\partial x^\nu}$$

Finally, we note that

$$\frac{\partial \sqrt{-g}}{\partial x^\nu} = \frac{-1}{2\sqrt{-g}} \frac{\partial g}{\partial x^\nu}$$

Hence,

$$\frac{1}{\sqrt{-g}} \frac{\partial \sqrt{-g}}{\partial x^\nu} = \frac{1}{2g} \frac{\partial g}{\partial x^\nu}$$

so that

$$\Gamma^{\mu}_{\nu\mu} = \frac{1}{\sqrt{-g}} \frac{\partial \sqrt{-g}}{\partial x^\nu}$$

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Remark: Another proof of $dg = g g^{\mu\nu} dg_{\mu\nu}$

From the definition of the determinant, we have

$$g_{\mu\rho} g_{\nu\sigma} g_{\alpha\delta} g_{\beta\tau} \epsilon^{\rho\sigma\delta\tau} = -g \epsilon_{\mu\nu\alpha\beta}$$

Note the minus sign, since by convention $\epsilon^{0123} = -\epsilon_{0123} = 1$.

Take the differential of this equation to obtain:

$$-dg \epsilon_{\mu\nu\alpha\beta} = \epsilon^{\rho\sigma\delta\tau} \left[dg_{\mu\rho} g_{\nu\sigma} g_{\alpha\delta} g_{\beta\tau} + g_{\mu\rho} dg_{\nu\sigma} g_{\alpha\delta} g_{\beta\tau} + g_{\mu\rho} g_{\nu\sigma} dg_{\alpha\delta} g_{\beta\tau} + g_{\mu\rho} g_{\nu\sigma} g_{\alpha\delta} dg_{\beta\tau} \right]$$

We also note that

$$g_{\nu\sigma} g_{\alpha\delta} g_{\beta\tau} \epsilon^{\rho\sigma\delta\tau} = -g g^{\rho\delta} \epsilon_{\nu\alpha\beta\rho}$$

since $g^{\mu\rho}$ is the inverse of $g_{\mu\rho}$. Likewise

$$g_{\mu\rho} g_{\alpha\delta} g_{\beta\tau} \epsilon^{\rho\sigma\delta\tau} = -g g^{\delta\sigma} \epsilon_{\mu\nu\alpha\beta}$$

$$g_{\mu\rho} g_{\nu\sigma} g_{\beta\tau} \epsilon^{\rho\sigma\delta\tau} = -g g^{\delta\sigma} \epsilon_{\mu\nu\alpha\beta}$$

$$g_{\mu\rho} g_{\nu\sigma} g_{\alpha\delta} \epsilon^{\rho\sigma\delta\tau} = -g g^{\delta\tau} \epsilon_{\mu\nu\alpha\beta}$$

Thus,

$$dg \epsilon_{\mu\nu\alpha\beta} = g \left[g^{\rho\delta} dg_{\mu\rho} \epsilon_{\nu\alpha\beta\delta} + g^{\delta\sigma} dg_{\nu\sigma} \epsilon_{\mu\alpha\beta\delta} + g^{\delta\sigma} dg_{\alpha\delta} \epsilon_{\mu\nu\beta\sigma} + g^{\delta\tau} dg_{\beta\tau} \epsilon_{\mu\nu\alpha\delta} \right]$$

Multiply by $\epsilon^{\mu\nu\alpha\beta}$ and sum over repeated indices.

Using:

$$\epsilon^{\mu\nu\alpha\beta} \epsilon_{\mu\nu\alpha\beta} = -24$$

$$\epsilon^{\mu\nu\alpha\beta} \epsilon_{\delta\nu\alpha\beta} = -6 \delta^{\mu}_{\delta}$$

etc., we obtain:

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$$dg = \frac{g}{24} 6 \left[g^{\mu\rho} dg_{\mu\rho} + g^{\nu\sigma} dg_{\nu\sigma} + g^{\alpha\delta} dg_{\alpha\delta} + g^{\beta\tau} dg_{\beta\tau} \right]$$

All four terms are identical after relabeling. Thus, we end up with

$$dg = g g^{\mu\nu} dg_{\mu\nu}$$

as before. We now turn to (a) and (b).

(a) Start with $A^{\mu}_{;\nu} = A^{\mu}_{,\nu} + \Gamma^{\mu}_{\rho\nu} A^{\rho}$. Set $\nu = \mu$ and sum over μ . Using $\Gamma^{\mu}_{\rho\mu} = \Gamma^{\mu}_{\mu\rho}$, we can write:

$$A^{\mu}_{;\mu} = A^{\mu}_{,\mu} + \Gamma^{\mu}_{\mu\beta} A^{\beta} = A^{\mu}_{,\mu} + \frac{1}{\sqrt{-g}} \frac{\partial \sqrt{-g}}{\partial x^{\beta}} A^{\beta}$$

This should be compared with:

$$\frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^{\mu}} (\sqrt{-g} A^{\mu}) = \frac{\partial}{\partial x^{\mu}} A^{\mu} + A^{\mu} \frac{1}{\sqrt{-g}} \frac{\partial \sqrt{-g}}{\partial x^{\mu}}$$

and we see that indeed

$$A^{\mu}_{;\mu} = \frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^{\mu}} (\sqrt{-g} A^{\mu})$$

(b) Start with $F^{\mu\nu}_{;\alpha} = F^{\mu\nu}_{,\alpha} + \Gamma^{\mu}_{\beta\alpha} F^{\beta\nu} + \Gamma^{\nu}_{\beta\alpha} F^{\mu\beta}$. Set $\alpha = \mu$ and sum over μ . It follows that:

$$F^{\mu\nu}_{;\mu} = F^{\mu\nu}_{,\mu} + \Gamma^{\mu}_{\mu\beta} F^{\beta\nu} + \Gamma^{\nu}_{\mu\beta} F^{\mu\beta}$$

By assumption, $F^{\mu\nu}$ is an antisymmetric tensor. Therefore, $\Gamma^{\nu}_{\mu\beta} F^{\mu\beta} = 0$ since $F^{\mu\beta}$ is antisymmetric whereas $\Gamma^{\nu}_{\mu\beta}$ is symmetric under the interchange of μ and β . Hence,

$$F^{\mu\nu}_{;\mu} = F^{\mu\nu}_{,\mu} + \frac{1}{\sqrt{-g}} \frac{\partial \sqrt{-g}}{\partial x^{\beta}} F^{\beta\nu}$$

The ν index goes along for the ride. Otherwise, the result is similar to that of part (a), so that

$$F^{\mu\nu}_{;\mu} = \frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^{\mu}} (\sqrt{-g} F^{\mu\nu}) \quad \text{if } F^{\mu\nu} = -F^{\nu\mu}$$

(9)

$$\textcircled{4} \quad ds^2 = \left(1 - \frac{2GM}{c^2 r}\right) c^2 dt^2 - \left(1 - \frac{2GM}{c^2 r}\right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

(a) Consider the Lagrangian

$$L = \left(1 - \frac{2GM}{c^2 r}\right) (\dot{x}^0)^2 - \left(1 - \frac{2GM}{c^2 r}\right)^{-1} \dot{r}^2 - r^2 \dot{\theta}^2 - r^2 \sin^2 \theta \dot{\phi}^2$$

Using the Lagrange equations:

$$\frac{d}{ds} \left(\frac{\partial L}{\partial \dot{x}^0} \right) = \frac{\partial L}{\partial x^0} \implies \frac{d}{ds} \left[2\dot{x}^0 \left(1 - \frac{2GM}{c^2 r}\right) \right] = 0$$

$$\text{or } \ddot{x}^0 \left(1 - \frac{2GM}{c^2 r}\right) + \frac{2GM}{c^2 r^2} \dot{r} \dot{x}^0 = 0$$

which we rewrite as

$$\ddot{x}^0 + \frac{2GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1} \dot{r} \dot{x}^0 = 0$$

$$\frac{d}{ds} \left(\frac{\partial L}{\partial \dot{r}} \right) = \frac{\partial L}{\partial r} \implies$$

$$\frac{d}{ds} \left[-2\dot{r} \left(1 - \frac{2GM}{c^2 r}\right)^{-1} \right] = \frac{2GM}{c^2 r^2} (\dot{x}^0)^2 + \left(1 - \frac{2GM}{c^2 r}\right)^{-2} \left(\frac{2GM}{c^2 r^2} \right) \dot{r}^2 - 2r\dot{\theta}^2 - 2r\sin^2 \theta \dot{\phi}^2$$

Hence,

$$\ddot{r} \left(1 - \frac{2GM}{c^2 r}\right)^{-1} - \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-2} \dot{r}^2 + \frac{GM}{c^2 r^2} (\dot{x}^0)^2 - r\dot{\theta}^2 - r\sin^2 \theta \dot{\phi}^2 = 0$$

Multiply by $1 - \frac{2GM}{c^2 r}$ to obtain:

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$$\ddot{r} - \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1} \dot{r}^2 + \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right) (\dot{x}^0)^2 - r \left(1 - \frac{2GM}{c^2 r}\right) \dot{\theta}^2 - r \sin^2 \theta \left(1 - \frac{2GM}{c^2 r}\right) \dot{\phi}^2 = 0$$

$$\frac{d}{ds} \left(\frac{\partial L}{\partial \dot{\theta}} \right) = \frac{\partial L}{\partial \theta} \implies \frac{d}{ds} [-2r^2 \dot{\theta}] = -2r^2 \sin \theta \cos \theta \dot{\phi}^2$$

That is,

$$\ddot{\theta} + \frac{2}{r} \dot{r} \dot{\theta} - \sin \theta \cos \theta \dot{\phi}^2 = 0$$

$$\frac{d}{ds} \left(\frac{\partial L}{\partial \dot{\phi}} \right) = \frac{\partial L}{\partial \phi} \implies \frac{d}{ds} [2r^2 \sin^2 \theta \dot{\phi}] = 0$$

$$\text{or } \sin^2 \theta \ddot{\phi} + \frac{2\dot{r}\dot{\phi}}{r} \sin^2 \theta + 2\sin \theta \cos \theta \dot{\theta} \dot{\phi} = 0$$

Divide by $\sin^2 \theta$ to obtain

$$\ddot{\phi} + \frac{2\dot{r}\dot{\phi}}{r} + 2\cot \theta \dot{\theta} \dot{\phi} = 0$$

We now compare with

$$\ddot{x}^\mu + \Gamma_{\alpha\beta}^\mu \dot{x}^\alpha \dot{x}^\beta = 0$$

and read off all the non-zero $\Gamma_{\alpha\beta}^\mu$. Keep in mind that $\Gamma_{\alpha\beta}^\mu = \Gamma_{\beta\alpha}^\mu$. Thus, for $\alpha \neq \beta$,

$$\Gamma_{\alpha\beta}^\mu \dot{x}^\alpha \dot{x}^\beta = \frac{1}{2} \Gamma_{\alpha\beta}^\mu (\dot{x}^\alpha \dot{x}^\beta + \dot{x}^\beta \dot{x}^\alpha)$$

(11)

The result:

$$\Gamma_{or}^o = \Gamma_{ro}^o = \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{rr}^r = -\frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{oo}^r = \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{\theta\theta}^r = -r \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{\phi\phi}^r = -r \sin^2 \theta \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{r\theta}^\theta = \Gamma_{\theta r}^\theta = \frac{1}{r}$$

$$\Gamma_{\phi\phi}^\theta = -\sin \theta \cos \theta$$

$$\Gamma_{\phi r}^\phi = \Gamma_{r\phi}^\phi = \frac{1}{r}$$

$$\Gamma_{\theta\phi}^\phi = \Gamma_{\phi\theta}^\phi = \cot \theta$$

These are the thirteen non-zero Christoffel symbols.

(b) Use $\Gamma_{\alpha\beta}^\mu = \frac{1}{2} g^{\mu\nu} [g_{\nu\beta,\alpha} + g_{\nu\alpha,\beta} - g_{\alpha\beta,\nu}]$

$$\begin{aligned} \text{Thus, } \Gamma_{or}^o &= \frac{1}{2} g^{oo} \left[\frac{\partial g_{or}}{\partial x^o} + \frac{\partial g_{oo}}{\partial r} - \frac{\partial g_{or}}{\partial x^o} \right] \\ &= \frac{1}{2} g^{oo} \frac{\partial g_{oo}}{\partial r} \end{aligned}$$

Since $g_{\mu\nu}$ is diagonal, we have $g^{\mu\mu} = \frac{1}{g_{\mu\mu}}$ (no sum over μ).

Thus, $g^{oo} = \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$. Next,

$$\frac{\partial g_{oo}}{\partial r} = \frac{2GM}{c^2 r^2}$$

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Thus, $\Gamma_{or}^o = \Gamma_{ro}^o = \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$

The other non-zero Christoffel symbols are also easy to obtain:

$$\Gamma_{oo}^r = -\frac{1}{2} g^{rr} \frac{\partial g_{oo}}{\partial r} = \frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{rr}^r = \frac{1}{2} g^{rr} \frac{\partial g_{rr}}{\partial r} = -\frac{GM}{c^2 r^2} \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{\theta\theta}^r = -\frac{1}{2} g^{rr} \frac{\partial g_{\theta\theta}}{\partial r} = -r \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{\phi\phi}^r = -\frac{1}{2} g^{rr} \frac{\partial g_{\phi\phi}}{\partial r} = -r \sin^2 \theta \left(1 - \frac{2GM}{c^2 r}\right)^{-1}$$

$$\Gamma_{r\theta}^\theta = \Gamma_{\theta r}^\theta = \frac{1}{2} g^{\theta\theta} \frac{\partial g_{\theta\theta}}{\partial r} = \frac{1}{2} \left(-\frac{1}{r^2}\right) (-2r) = \frac{1}{r}$$

$$\Gamma_{r\phi}^\phi = \Gamma_{\phi r}^\phi = \frac{1}{2} g^{\phi\phi} \frac{\partial g_{\phi\phi}}{\partial r} = \frac{1}{2} \left(-\frac{1}{r^2 \sin^2 \theta}\right) (-2r \sin^2 \theta) = \frac{1}{r}$$

$$\Gamma_{\phi\phi}^\theta = -\frac{1}{2} g^{\theta\theta} \frac{\partial g_{\phi\phi}}{\partial \theta} = -\frac{1}{2r^2} 2r^2 \sin \theta \cos \theta = -\sin \theta \cos \theta$$

$$\begin{aligned} \Gamma_{\theta\phi}^\phi &= \Gamma_{\phi\theta}^\phi = \frac{1}{2} g^{\phi\phi} \frac{\partial g_{\phi\phi}}{\partial \theta} = \frac{1}{2r^2 \sin^2 \theta} 2r^2 \sin \theta \cos \theta \\ &= \cot \theta \end{aligned}$$

All other Γ 's vanish.

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⑤ The metric for the surface of the sphere of radius one is
 $ds^2 = d\theta^2 + \sin^2\theta d\phi^2$.

We obtain the geodesic equations from $L = \dot{\theta}^2 + \sin^2\theta \dot{\phi}^2$.

$$\frac{d}{ds} \left(\frac{\partial L}{\partial \dot{\theta}} \right) = \frac{\partial L}{\partial \theta} \implies \frac{d}{ds} (2\dot{\theta}) = 2\sin\theta \cos\theta \dot{\phi}^2$$

or
$$\ddot{\theta} - \sin\theta \cos\theta \dot{\phi}^2 = 0$$

$$\frac{d}{ds} \left(\frac{\partial L}{\partial \dot{\phi}} \right) = \frac{\partial L}{\partial \phi} \implies \frac{d}{ds} (2\dot{\phi} \sin^2\theta) = 0$$

or
$$\dot{\phi} \sin^2\theta + 2\sin\theta \cos\theta \dot{\theta} \dot{\phi} = 0$$

$$\ddot{\phi} + 2\cot\theta \dot{\theta} \dot{\phi} = 0$$

Lines of longitude on the sphere correspond to

$$\phi = \text{constant}$$

$$\theta = s$$

where s is the arclength measured along the geodesic. It is trivial to check that this satisfies the two geodesic equations given above.

Remark: It is not much more difficult to obtain all geodesics. From the equation

$$\frac{d}{ds} (\dot{\phi} \sin^2\theta) = 0$$

we learn that $\dot{\phi} \sin^2\theta = J$ where J is a constant. Insert this result into the first geodesic equation and eliminate $\dot{\phi}$. But, there is an easier method. Noting that

$$1 = \left(\frac{d\theta}{ds} \right)^2 + \sin^2\theta \left(\frac{d\phi}{ds} \right)^2$$

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we see that the value of L on the geodesic is $L=1$. Thus,

$$\dot{\theta}^2 + \sin^2\theta \dot{\phi}^2 = 1$$

Putting $\dot{\phi} \sin^2\theta = J$, we have

$$\dot{\theta}^2 = 1 - \frac{J^2}{\sin^2\theta}$$

$$\frac{d\theta}{ds} = \frac{\sqrt{\sin^2\theta - J^2}}{\sin\theta}$$

Define $J = \sin\theta_0$. Then, we can integrate this to obtain

$$\cos\theta = \cos\theta_0 \sin(s-s_0)$$

where s_0 is the constant of integration. For simplicity, set $s_0 = 0$ which simply defines the origin of the arclength variable.

$$\cos\theta = \cos\theta_0 \sin s$$

In this case, $s=0$ corresponds to $\theta = \frac{\pi}{2}$ (the equator). Finally, from

$$\dot{\phi} \sin^2\theta = \sin\theta_0$$

thus,

$$\frac{d\phi}{ds} = \frac{\sin\theta_0}{1 - \cos^2\theta_0 \sin^2 s}$$

Integration gives:

$$\tan(\phi - \phi_0) = \sin\theta_0 \tan s$$

where ϕ_0 is the constant of integration. The two boxed equations provide a parametric representation of any geodesic, and depend on the constants θ_0 and ϕ_0 . For example, if $\theta_0 = 0$, then we have

$$\phi = \phi_0$$

$$\theta = \frac{\pi}{2} - s$$

which describes one of the longitudes. If $\theta_0 = \frac{\pi}{2}$, then we obtain

$$\theta = \frac{\pi}{2}$$

$$\phi = \phi_0 + s$$

which is the equator.

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Addendum to Problem 3

The key result

$$\frac{1}{g} \frac{\partial g}{\partial x^B} = g^{\mu\nu} \frac{\partial g_{\mu\nu}}{\partial x^B}$$

can be proved in other ways. Here are two alternate methods:

Method 1: For any matrix A ,

$$\det(\exp A) = \exp(\operatorname{tr} A)$$

where

$$\exp A \equiv \sum_{n=0}^{\infty} \frac{A^n}{n!}$$

is defined by the usual power series of the exponential (which converges for all matrices). Take $\frac{\partial}{\partial x^B}$ of both sides

$$\frac{\partial}{\partial x^B} \det(\exp A) = \frac{\partial}{\partial x^B} \exp(\operatorname{tr} A)$$

assuming that the matrix elements of A depend on x^B . Now,

$$\begin{aligned} \frac{\partial}{\partial x^B} \exp(\operatorname{tr} A) &= \exp(\operatorname{tr} A) \frac{\partial}{\partial x^B} \operatorname{tr} A \\ &= \exp(\operatorname{tr} A) \operatorname{tr} \left(\frac{\partial A}{\partial x^B} \right) \end{aligned}$$

where we have used the fact that the trace depends linearly on the matrix elements (since it is the sum of the diagonal elements), and the derivative is a linear operator. At this point, note that

$$\frac{\partial}{\partial x^B} \exp A = \exp A \frac{\partial A}{\partial x^B}$$

(which can be proved using the chain rule and the power series definition of $\exp A$). Thus,

$$\frac{\partial A}{\partial x^B} = (\exp A)^{-1} \frac{\partial}{\partial x^B} \exp A$$

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Thus, if $B = \exp A$, we have proven that

$$\frac{\partial}{\partial x^B} \det B = \det B \operatorname{tr} \left(B^{-1} \frac{\partial B}{\partial x^B} \right)$$

Putting $B = g_{\mu\nu}$, $B^{-1} = g^{\mu\nu}$ and $\det B = g$ yields the desired result.Remarks: If $B = \exp A$, then we can write $A = \ln B$.

In this case, the above proof can be rewritten as follows:

$$\ln(\det B) = \operatorname{tr}(\ln B)$$

$$\frac{\partial}{\partial x^B} \ln(\det B) = \frac{1}{\det B} \frac{\partial}{\partial x^B} \det B$$

$$\frac{\partial}{\partial x^B} \operatorname{tr}(\ln B) = \operatorname{tr} \left(\frac{\partial}{\partial x^B} \ln B \right) = \operatorname{tr} \left(B^{-1} \frac{\partial B}{\partial x^B} \right)$$

from which it follows that

$$\frac{1}{\det B} \frac{\partial}{\partial x^B} \det B = \operatorname{tr} \left(B^{-1} \frac{\partial B}{\partial x^B} \right)$$

[Note that the proof of

$$\frac{\partial}{\partial x^B} \ln B = B^{-1} \frac{\partial B}{\partial x^B}$$

is easy if $B = \exp A$, since $\ln(\exp A) = A$ by definition.]The only remaining question is whether we can apply these results to $g_{\mu\nu}$. Note the implicit assumption that there exists an A such that $B = \exp A$ (or equivalently that $\ln B$ exists).Since $g_{\mu\nu}$ is a real symmetric matrix (and by assumption it is non-singular), it can be diagonalized by an orthogonal transformation. Let G be the matrix $g_{\mu\nu}$. Then

$$O^T G O = \operatorname{diag}(\lambda_1, \lambda_2, \lambda_3, \lambda_4)$$

where diag is a matrix with diagonal elements given by the eigenvalues.

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Then,

$$O^T \ln G O = \text{diag}(\ln \lambda_1, \ln \lambda_2, \ln \lambda_3, \ln \lambda_4)$$

and

$$\ln G = O \text{diag}(\ln \lambda_1, \ln \lambda_2, \ln \lambda_3, \ln \lambda_4) O^T$$

If $g_{\mu\nu}$ is positive definite, then all the $\lambda_i > 0$ and $\ln G$ is well defined

However, in general relativity, $g_{\mu\nu}$ has one positive eigenvalue and three negative eigenvalues. We can still define $\ln G$, but we must be careful since the log of a negative number is multivalued.

Nevertheless, the end result,

$$\frac{1}{\det B} \frac{\partial}{\partial x^A} \det B = \text{tr} \left(B^{-1} \frac{\partial B}{\partial x^A} \right)$$

is true for all matrices B , and does not depend on ambiguities in the meaning of $\ln B$.

By the way, in the solution to problem 3, I derived

$$\frac{1}{\det A} d(\det A) = \text{Tr}(A^{-1} dA)$$

If I integrate this equation, then I obtain

$$\int \frac{1}{\det A} d(\det A) = \ln(\det A)$$

and

$$\int \text{Tr}(A^{-1} dA) = \text{Tr} \int A^{-1} dA = \text{Tr}(\ln A)$$

In the last line, I used the linear dependence of the trace on the matrix elements and the linearity of the integral. The final equality is formal but suggestive, and it follows that

$$\ln(\det A) = \text{Tr}(\ln A)$$

There is no integration constant, since this result is trivially satisfied when $A = 1$ (the identity matrix). A more rigorous proof would replace A above with $\exp A$. It is easy to see how to obtain:

$$\det(\exp A) = \exp(\text{tr} A)$$

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Method 2: In this method, we compute $\frac{\partial}{\partial x^A} \det B$ more directly.

First, let us review some matrix algebra. For any nonsingular matrix B , its inverse is given by

$$B^{-1} = \frac{1}{\det B} B^{\text{adj}}$$

where B^{adj} is the "adjoint" of B (not to be confused with hermitian adjoint B^\dagger) which is defined by:

$$B^{\text{adj}}_{ji} = \text{cof}(B_{ij})$$

(note the order of indices), where the cofactor of B_{ij} is given by

$$\text{cof}(B_{ij}) \equiv (-1)^{i+j} \det \begin{pmatrix} B_{11} & \dots & B_{1,j-1} & B_{1,j+1} & \dots & B_{1n} \\ \vdots & & \vdots & \vdots & & \vdots \\ B_{i-1,1} & \dots & B_{i-1,j-1} & B_{i-1,j+1} & \dots & B_{i-1,n} \\ B_{i+1,1} & \dots & B_{i+1,j-1} & B_{i+1,j+1} & \dots & B_{i+1,n} \\ \vdots & & \vdots & \vdots & & \vdots \\ B_{n1} & \dots & B_{n,j-1} & B_{n,j+1} & \dots & B_{nn} \end{pmatrix}$$

That is, the cofactor of B_{ij} is obtained by deleting the i^{th} row and j^{th} column from B , taking the determinant of the resulting $(n-1) \times (n-1)$ matrix and multiplying by a sign $(-1)^{i+j}$.

One can compute $\det B$ using the so-called cofactor expansion (or expansion by minors). One has the following formula:

$$\det B = \sum_{j=1}^n B_{ij} \text{cof}(B_{ij})$$

for fixed i (any choice of i gives the same result).

If B_{ij} depends on x , then

$$\frac{\partial}{\partial B_{ij}} \det B = \text{cof}(B_{ij})$$

since $\text{cof}(B_{ij})$ does not depend on B_{ij} since this element has been deleted from the matrix in computing the determinant (see above).

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By the chain rule,

$$\begin{aligned}
 \frac{\partial}{\partial x^{\beta}} \det B &= \frac{\partial B_{ij}}{\partial x^{\beta}} \frac{\partial}{\partial B_{ij}} \det B \\
 &= \frac{\partial B_{ij}}{\partial x^{\beta}} \operatorname{cof}(B_{ij}) \\
 &= \frac{\partial B_{ij}}{\partial x^{\beta}} B_{ji}^{\operatorname{adj}} \\
 &= \operatorname{Tr} \left(B^{\operatorname{adj}} \frac{\partial B}{\partial x^{\beta}} \right) \\
 &= \det B \operatorname{Tr} \left(B^{-1} \frac{\partial B}{\partial x^{\beta}} \right)
 \end{aligned}$$

where we have used $B^{-1} = \frac{1}{\det B} B^{\operatorname{adj}}$. The factor $\det B$ is just a number which we can pull out of the trace. Thus,

$$\frac{1}{\det B} \frac{\partial}{\partial x^{\beta}} \det B = \operatorname{Tr} \left(B^{-1} \frac{\partial B}{\partial x^{\beta}} \right)$$

for any non-singular matrix B .
