

$$\begin{aligned}
 \int \frac{d^n q}{(2\pi)^n} \frac{1}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 2)}{\Gamma(r)} \\
 \int \frac{d^n q}{(2\pi)^n} \frac{q^\mu}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= -i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 2)}{\Gamma(r)} p^\mu \\
 \int \frac{d^n q}{(2\pi)^n} \frac{q^\mu q^\nu}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 3)}{\Gamma(r)} \\
 &\quad \times [(\epsilon + r - 3)p^\mu p^\nu - \frac{1}{2}g^{\mu\nu}(p^2 + m^2)] \\
 \int \frac{d^n q}{(2\pi)^n} \frac{q^\mu q^\nu q^\alpha}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= -i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 3)}{\Gamma(r)} \\
 &\quad \times [(\epsilon + r - 3)p^\mu p^\nu p^\alpha - \frac{1}{2}(g^{\mu\nu}p^\alpha + g^{\mu\alpha}p^\nu + g^{\nu\alpha}p^\mu)(p^2 + m^2)] \\
 \int \frac{d^n q}{(2\pi)^n} \frac{q^\mu q^\nu q^\alpha q^\beta}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 4)}{\Gamma(r)} \\
 &\quad \times \left\{ (\epsilon + r - 3)(\epsilon + r - 4)p^\mu p^\nu p^\alpha p^\beta - \frac{1}{2}(\epsilon + r - 4)(g^{\mu\nu}p^\alpha p^\beta \right. \\
 &\quad \left. + g^{\mu\alpha}p^\nu p^\beta + g^{\mu\beta}p^\mu p^\alpha + g^{\nu\alpha}p^\mu p^\beta + g^{\nu\beta}p^\mu p^\alpha + g^{\alpha\beta}p^\mu p^\nu)(p^2 + m^2) \right. \\
 &\quad \left. + \frac{1}{4}(g^{\mu\nu}g^{\alpha\beta} + g^{\mu\alpha}g^{\nu\beta} + g^{\mu\beta}g^{\nu\alpha})(p^2 + m^2)^2 \right\}
 \end{aligned}$$

where ε (not to be confused with ϵ) is a positive infinitesimal constant and ϵ is related to the number of spacetime dimensions via

$$g^{\mu\nu}g_{\mu\nu} = n \equiv 4 - 2\epsilon.$$

In addition, we shall *define*

$$\int \frac{d^n q}{(2\pi)^n} \frac{1}{(q^2)^r} \equiv 0,$$

which corresponds to setting $p^2 = m^2 = 0$ in the first integral above under the assumption that $\epsilon < 2 - r$. However, in the dimensional regularization procedure, we shall adopt the above definition for *all* r .

We can expand about $\epsilon = 0$ by using

$$\Gamma(-N + \epsilon) = \frac{(-1)^N}{N!} \left[\frac{1}{\epsilon} + \psi(N + 1) + \mathcal{O}(\epsilon) \right],$$

where N is a non-negative integer, $\psi(x) \equiv \Gamma'(x)/\Gamma(x)$ with $\Gamma'(x) \equiv d\Gamma(x)/dx$,

$$\psi(1) = -\gamma, \quad \psi(N+1) = -\gamma + \sum_{k=1}^N \frac{1}{k},$$

and $\gamma = -\Gamma'(1) = 0.5772\cdots$ is the Euler-Mascheroni constant. If the $\mathcal{O}(\epsilon)$ terms are needed, then one must use $x\Gamma(x) = \Gamma(x+1)$ until $\Gamma(1+\epsilon)$ is reached, and then use

$$\log \Gamma(1+\epsilon) = -\gamma\epsilon + \sum_{k=2}^{\infty} \frac{(-\epsilon)^k}{k} \zeta(k),$$

where $\zeta(k)$ is the Riemann zeta function.

When fermions are involved, we need to consider the Dirac matrix algebra in n -dimensions. The Dirac gamma matrices are denoted by $\gamma^0, \gamma^1, \gamma^2, \dots, \gamma^{n-1}$. The n -dimensional analog of γ_5 is somewhat problematical, although we shall treat it as anticommuting with all the other gamma matrices. The following n -dimensional gamma matrix relations must be used:

1. $\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$
2. $\{\gamma^\mu, \gamma_5\} = 0$
3. $(\gamma_5)^2 = \mathbb{1}$
4. $\gamma_\mu \gamma^\mu = n$
5. $\gamma_\mu \gamma^\alpha \gamma^\mu = (2-n)\gamma^\alpha$
6. $\gamma_\mu \gamma^\alpha \gamma^\beta \gamma^\mu = 4g^{\alpha\beta} + (n-4)\gamma^\alpha \gamma^\beta$
7. $\gamma_\mu \gamma^\alpha \gamma^\beta \gamma^\rho \gamma^\mu = -2\gamma^\rho \gamma^\beta \gamma^\alpha + (4-n)\gamma^\alpha \gamma^\beta \gamma^\rho$
8. Trace formulae are unchanged. In particular, $\text{Tr}\gamma^\mu \gamma^\nu = 4g^{\mu\nu}$.

The 4 in the last trace formula is purely conventional. However, note that it is crucial to use $g^{\mu\nu} g_{\mu\nu} = n$ in all calculations in n -dimensions before taking the $\epsilon \rightarrow 0$ limit.

Finally, we record some of the Feynman parameter formulae:

$$\begin{aligned} \frac{1}{A^\alpha B^\beta} &= \frac{\Gamma(\alpha + \beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 dx \frac{x^{\alpha-1}(1-x)^{\beta-1}}{[xA + (1-x)B]^{\alpha+\beta}} \\ \frac{1}{A^\alpha B^\beta C^\delta} &= \frac{\Gamma(\alpha + \beta + \delta)}{\Gamma(\alpha)\Gamma(\beta)\Gamma(\delta)} \int_0^1 x dx \int_0^1 dy \frac{x^{\alpha+\beta-2} y^{\alpha-1}(1-x)^{\delta-1}(1-y)^{\beta-1}}{[xyA + x(1-y)B + (1-x)C]^{\alpha+\beta+\delta}} \end{aligned}$$

and more generally,

$$\begin{aligned} \frac{1}{A_1^{\alpha_1} A_2^{\alpha_2} \cdots A_N^{\alpha_N}} &= \frac{\Gamma(\alpha_1 + \alpha_2 + \cdots + \alpha_N)}{\Gamma(\alpha_1)\Gamma(\alpha_2)\cdots\Gamma(\alpha_N)} \int_0^1 dx_1 \int_0^1 dx_2 \cdots \int_0^1 dx_N \delta \left(\sum_{j=1}^N x_j - 1 \right) \\ &\quad \times \frac{x_1^{\alpha_1-1} x_2^{\alpha_2-1} \cdots x_N^{\alpha_N-1}}{(x_1 A_1 + x_2 A_2 + \cdots + x_N A_N)^{\alpha_1+\alpha_2+\cdots+\alpha_N}} \end{aligned}$$