

$$\begin{aligned}
\int \frac{d^n q}{(2\pi)^n} \frac{1}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 2)}{\Gamma(r)} \\
\int \frac{d^n q}{(2\pi)^n} \frac{q^\mu}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= -i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 2)}{\Gamma(r)} p^\mu \\
\int \frac{d^n q}{(2\pi)^n} \frac{q^\mu q^\nu}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 3)}{\Gamma(r)} \\
&\quad \times [(\epsilon + r - 3)p^\mu p^\nu - \tfrac{1}{2}g^{\mu\nu}(p^2 + m^2)] \\
\int \frac{d^n q}{(2\pi)^n} \frac{q^\mu q^\nu q^\alpha}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= -i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 3)}{\Gamma(r)} \\
&\quad \times [(\epsilon + r - 3)p^\mu p^\nu p^\alpha - \tfrac{1}{2}(g^{\mu\nu}p^\alpha + g^{\mu\alpha}p^\nu + g^{\nu\alpha}p^\mu)(p^2 + m^2)] \\
\int \frac{d^n q}{(2\pi)^n} \frac{q^\mu q^\nu q^\alpha q^\beta}{(q^2 + 2q \cdot p - m^2 + i\varepsilon)^r} &= i(-1)^r (p^2 + m^2)^{2-\epsilon-r} (4\pi)^{\epsilon-2} \frac{\Gamma(\epsilon + r - 4)}{\Gamma(r)} \\
&\quad \times \left\{ (\epsilon + r - 3)(\epsilon + r - 4)p^\mu p^\nu p^\alpha p^\beta - \tfrac{1}{2}(\epsilon + r - 4)(g^{\mu\nu}p^\alpha p^\beta \right. \\
&\quad + g^{\mu\alpha}p^\nu p^\beta + g^{\mu\beta}p^\nu p^\alpha + g^{\nu\alpha}p^\mu p^\beta + g^{\nu\beta}p^\mu p^\alpha + g^{\alpha\beta}p^\mu p^\nu)(p^2 + m^2) \\
&\quad \left. + \tfrac{1}{4}(g^{\mu\nu}g^{\alpha\beta} + g^{\mu\alpha}g^{\nu\beta} + g^{\mu\beta}g^{\nu\alpha})(p^2 + m^2)^2 \right\}
\end{aligned}$$

where  $\varepsilon$  (not to be confused with  $\epsilon$ ) is a positive infinitesimal constant and  $\epsilon$  is related to the number of spacetime dimensions via

$$g^{\mu\nu}g_{\mu\nu} = n \equiv 4 - 2\epsilon.$$

In addition, we shall *define*

$$\int \frac{d^n q}{(2\pi)^n} \frac{1}{(q^2)^r} \equiv 0,$$

which corresponds to setting  $p^2 = m^2 = 0$  in the first integral above under the assumption that  $\epsilon < 2 - r$ . However, in the dimensional regularization procedure, we shall adopt the above definition for *all*  $r$ .

We can expand about  $\epsilon = 0$  by using

$$\Gamma(-N + \epsilon) = \frac{(-1)^N}{N!} \left[ \frac{1}{\epsilon} + \psi(N + 1) + \mathcal{O}(\epsilon) \right],$$

where  $N$  is a non-negative integer,  $\psi(x) \equiv \Gamma'(x)/\Gamma(x)$  with  $\Gamma'(x) \equiv d\Gamma(x)/dx$ ,

$$\psi(1) = -\gamma, \quad \psi(N+1) = -\gamma + \sum_{k=1}^N \frac{1}{k},$$

and  $\gamma = -\Gamma'(1) = 0.5772 \dots$  is the Euler-Mascheroni constant. If the  $\mathcal{O}(\epsilon)$  terms are needed, then one must use  $x\Gamma(x) = \Gamma(x+1)$  until  $\Gamma(1+\epsilon)$  is reached, and then use

$$\log \Gamma(1+\epsilon) = -\gamma\epsilon + \sum_{k=2}^{\infty} \frac{(-\epsilon)^k}{k} \zeta(k),$$

where  $\zeta(k)$  is the Riemann zeta function.

When fermions are involved, we need to consider the Dirac matrix algebra in  $n$ -dimensions. The Dirac gamma matrices are denoted by  $\gamma^0, \gamma^1, \gamma^2, \dots, \gamma^{n-1}$ . The following  $n$ -dimensional gamma matrix relations must be used:

1.  $\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$
2.  $\gamma_\mu \gamma^\mu = n$
3.  $\gamma_\mu \gamma^\alpha \gamma^\mu = (2-n)\gamma^\alpha$
4.  $\gamma_\mu \gamma^\alpha \gamma^\beta \gamma^\mu = 4g^{\alpha\beta} + (n-4)\gamma^\alpha \gamma^\beta$
5.  $\gamma_\mu \gamma^\alpha \gamma^\beta \gamma^\rho \gamma^\mu = -2\gamma^\rho \gamma^\beta \gamma^\alpha + (4-n)\gamma^\alpha \gamma^\beta \gamma^\rho$
6.  $\gamma_\mu \gamma_\alpha \gamma_\beta \gamma_\rho \gamma_\sigma \gamma^\mu = 2(\gamma_\alpha \gamma_\sigma \gamma_\rho \gamma_\beta + \gamma_\beta \gamma_\rho \gamma_\sigma \gamma_\alpha) + (n-4)\gamma_\alpha \gamma_\beta \gamma_\rho \gamma_\sigma$
7. Trace formulae are unchanged. In particular,  $\text{Tr } \gamma^\mu \gamma^\nu = 4g^{\mu\nu}$ , where the 4 in the trace formula is purely conventional. However, is crucial to use  $g^{\mu\nu} g_{\mu\nu} = n$  in all calculations in  $n$ -dimensions before taking the  $\epsilon \rightarrow 0$  limit.

The  $n$ -dimensional analog of  $\gamma_5$ , which satisfies  $(\gamma_5)^2 = \mathbb{1}$  and  $\text{Tr } \gamma_5 = \text{Tr } \gamma_5 \gamma_\mu \gamma_\nu = 0$ , is problematical since  $\epsilon_{\alpha\beta\rho\sigma}$  is inherently four-dimensional. In some applications, one can take the anticommutation relation,  $\{\gamma^\mu, \gamma_5\} = 0$ , to be valid for  $n \neq 4$ , but some care is required.

Finally, we record some of the Feynman parameter formulae:

$$\frac{1}{A^\alpha B^\beta} = \frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \int_0^1 dx \frac{x^{\alpha-1}(1-x)^{\beta-1}}{[xA + (1-x)B]^{\alpha+\beta}}$$

$$\frac{1}{A^\alpha B^\beta C^\delta} = \frac{\Gamma(\alpha+\beta+\delta)}{\Gamma(\alpha)\Gamma(\beta)\Gamma(\delta)} \int_0^1 x dx \int_0^1 dy \frac{x^{\alpha+\beta-2} y^{\alpha-1} (1-x)^{\delta-1} (1-y)^{\beta-1}}{[xyA + x(1-y)B + (1-x)C]^{\alpha+\beta+\delta}}$$

and more generally,

$$\frac{1}{A_1^{\alpha_1} A_2^{\alpha_2} \dots A_N^{\alpha_N}} = \frac{\Gamma(\alpha_1 + \alpha_2 + \dots + \alpha_N)}{\Gamma(\alpha_1)\Gamma(\alpha_2)\dots\Gamma(\alpha_N)} \int_0^1 dx_1 \int_0^1 dx_2 \dots \int_0^1 dx_N \delta\left(\sum_{j=1}^N x_j - 1\right)$$

$$\times \frac{x_1^{\alpha_1-1} x_2^{\alpha_2-1} \dots x_N^{\alpha_N-1}}{(x_1 A_1 + x_2 A_2 + \dots + x_N A_N)^{\alpha_1 + \alpha_2 + \dots + \alpha_N}}$$